On density concomitants of the covariant curvature tensor in the two- and three-dimensional Riemann space

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1. Introduction. One of the basic questions of the theory of geometric objects is to determine the algebraic concomitants of a certain type for a given object.

In the case n=2 the scalar concomitants of the mixed tensor have been determined by Goląb [8] under the supposition that the functions in question are of closs C^1 . All the scalar concomitants of the mixed tensor in the n-dimensional space have been determined by Aczél and Hosszú [2], while those of the twice covariant tensor in the n-dimensional space were determined by Zajtz in [14].

By means of the analytic method Bieszk [3] has determined for the curvature tensor concomitants being either densities or tensors of second order in the two-dimensional space and linear concomitants being two-times covariant tensors in the three-dimensional space.

The same analytic method, reducing the system of functional equations to the system of differential equations of the first order, has been applied by Bieszk and Węgrzynowski in [5] and [6] to determine densities and vector concomitants of the antisymmetric tensor and linear concomitants of a tensor T_{ab}^{ν} in the two-dimensional space.

Scalar concomitants of a tensor $T^{\nu}_{\alpha\beta}$ without regularity assumptions in the two-dimensional space were determined by Węgrzynowski in paper [13].

A certain general and uniform method reducing the determination of the concomitants of geometric objects to the question of determining certain special subgroups of the general linear group GL_n was given by Zajtz and Siwek in [12].

In this paper we shall determine by the analytic method all density (scalar) concomitants of the covariant curvature tensor in the two- and three-dimensional Riemann space.

Another method of solving the above-mentioned problem will be given in a forthcoming paper by S. Topa.

2. Density (scalar) concomitants in the two-dimensional space V_2 . First of all we give some general notations. If the passage from one allowable coordinate system (λ) to another (λ ') is given by the system of functions

(1)
$$\xi^{\lambda'} = \varphi^{\lambda'}(\xi^{\lambda}), \quad \lambda = 1, 2, \ldots, n, \ \lambda' = 1', 2', \ldots, n',$$

where

$$A_{\lambda}^{\lambda'} = \frac{\partial \varphi^{\lambda'}(\xi^{\lambda})}{\partial \xi^{\lambda}},$$

$$J \stackrel{\mathrm{df}}{=} \det(A_{\lambda}^{\lambda'}) \neq 0,$$

then for the inverse transformation

$$\xi^{\lambda} = \psi^{\lambda}(\xi^{\lambda'}),$$

we introduce the notation

(5)
$$A_{\lambda'}^{\lambda} \stackrel{\mathrm{df}}{=} \frac{\partial \psi^{\lambda}(\xi^{\lambda'})}{\partial \xi^{\lambda'}}.$$

Between $A_{\lambda}^{\lambda'}$ and $A_{\lambda'}^{\lambda}$ the following relations occur:

$$A_{\lambda'}^{\lambda} = \frac{\operatorname{minor} A_{\lambda}^{\lambda'}}{J}.$$

In a Riemann space V_n the induced connexion is given by means of the Christoffel symbols of second kind

where $g^{a\beta}$ is the inverse tensor to the metric tensor $g_{a\beta}$.

The curvature tensor (Riemann-Christoffel tensor) is defined as follons:

(8)
$$R_{a\beta\gamma}^{\delta} \stackrel{\text{df}}{=} 2\partial_{[a} \{_{\beta]\gamma}^{\delta}\} + 2\{_{[a|\varrho|}^{\delta}\} \{_{\beta]\gamma}^{\varrho}\}.$$

The so-called covariant curvature tensor is defined by

(9)
$$R_{a\beta\gamma\delta} \stackrel{\mathrm{df}}{=} R_{a\beta\gamma}^{\varrho} g_{\delta\varrho}.$$

Tensor (9) has the well-known properties

$$egin{align} 1^{
m o} & R_{aeta\gamma\delta} = R_{\gamma\delta aeta}, \ & \ 2^{
m o} & R_{aeta\gamma\delta} = -R_{eta a\gamma\delta} = -R_{aeta\delta\gamma}. \ \end{pmatrix}$$

The number N of the so-called essential components of the tensor $R_{\alpha\beta\gamma\delta}$ in space V_n is defined by the formula

(11)
$$N = \frac{n^2(n^2-1)}{12},$$

(see [11], 90.9).

After having passed to the new coordinate system (λ') the tensor $R_{\alpha\beta\gamma\delta}$ has the components $R_{\alpha'\beta'\gamma'\delta'}$ connected with $R_{\alpha\beta\gamma\delta}$ by the formula

$$(12) R_{\alpha'\beta'\gamma'\delta'} = A^a_{\alpha'}A^{\beta}_{\beta'}A^{\gamma}_{\gamma'}A^{\delta}_{\delta'}R_{\alpha\beta\gamma\delta}.$$

Let us proceed to determine the density concomitants of $R_{\alpha\beta\gamma\delta}$ in the two-dimensional Riemann space.

For n=2, N=1, i.e. $R_{a\beta\gamma\delta}$ has only one essential component

$$(13) x \stackrel{\text{df}}{=} R_{1212}.$$

For the matrix A_{λ}^{λ} we introduce the shorter notation

Hence we have

$$(15) J = (p_1 p_4 - p_2 p_3)^{-1},$$

and according to [11] (formula 92.1)

$$x' \stackrel{\mathrm{df}}{=} R_{1'2'1'2'} = (p_1 p_4 - p_2 p_3)^2 x,$$

or more briefly

$$(16) x' = J^{-2}x.$$

We seek an algebraic concomitant H of $R_{a\beta\gamma\delta}$, which is a density of weight (-r).

For n=2, H is a function of x fulfilling the equation

(17)
$$H(x') = \varepsilon |J|^r H(x)$$

 \mathbf{or}

(18)
$$H[(p_1p_4-p_2p_3)^2x] = \varepsilon |J|^r H(x),$$

where

(19)
$$\varepsilon = \begin{cases} 1 & \text{for a Weyl density,} \\ \operatorname{sgn} J & \text{for an ordinary density.} \end{cases}$$

Assuming that H(x) is of class C^1 , we reduce equation (17) to an ordinary differential one. We differentiate (18) with respect to p_1 , p_2 , p_3 , p_4 and next substitute

$$(20) p_1 = p_4 = 1, p_2 = p_3 = 0.$$

After this operation we get four equations reducing to

$$(21) 2xH'(x) = -rH(x),$$

in which there is no intervention of ε .

We have to distinguish two cases, I and II.

I. Let us assume that r=0.

Ia. If x = 0, the problem is trivial, because $R_{1212} = 0$ in every coordinate system.

Ib. If $x \neq 0$, then H'(x) = 0; hence H(x) = const and in this case only arbitrary scalars could be concomitants of the tensor R_{1212} .

II. Let us assume that $r \neq 0$.

IIa. If x = 0, then we have again the trivial case.

IIb. If $x \neq 0$, then the general solution of (21) has the form

(22)
$$H = \begin{cases} C_1 |x|^{-r/2} & \text{for } x > 0, \\ C_2 |x|^{-r/2} & \text{for } x < 0, \end{cases}$$

where C_1 , C_2 are arbitrary constants different from zero (different or equal).

It is easy to prove that if $\varepsilon = 1$, H(x) defined by (22) fulfils (17), while for $\varepsilon = \operatorname{sgn} x$, formula (22) is no solution of (17) in the whole domain but for x > 0 only.

Now we can state

THEOREM 1. In the space V_2 the only scalar concomitants (of the class C^1) of the tensor $R_{a\beta\gamma\delta}$ are arbitrary scalars, while the only density concomitants of the weight (-r) are Weyl-densities of the form

(23)
$$H(x) = C|x|^{-r/2},$$

where the arbitrary constant C different from zero is given by

(24)
$$C = \begin{cases} C_1 & \text{for } x > 0, \\ C_2 & \text{for } x < 0. \end{cases}$$

3. Density (scalar) concomitants in the Riemann space V_3 . In accordance with (11) (Section 2) we have for n=3, N=6. We introduce the following shorter notations for the six essential components of the tensor $R_{\alpha\beta\gamma\delta}$:

(1)
$$x_1 = R_{1212}, \quad x_2 = R_{1313}, \quad x_3 = R_{2323}, \\ x_4 = R_{1213}, \quad x_5 = E_{1223}, \quad x_6 = R_{1323}$$

and the shorter ones for the elements of the matrix $A_{\lambda}^{\lambda'}$:

(2)
$$[a_{ij}] \stackrel{\text{df}}{=} [A_j^{i'}], \text{ where } i, j = 1, 2, 3;$$

$$(3) J = \det(a_{ii}) \neq 0.$$

In accordance with (6) (Section 2) and (2) we have

$$(4) \qquad [A_{i'}^{j}] = J^{-1} \cdot \begin{bmatrix} a_{22}a_{33} - a_{23}a_{32} & a_{23}a_{31} - a_{21}a_{33} & a_{21}a_{32} - a_{22}a_{31} \\ a_{13}a_{32} - a_{12}a_{33} & a_{11}a_{33} - a_{13}a_{31} & a_{12}a_{31} - a_{11}a_{32} \\ a_{12}a_{23} - a_{13}a_{22} & a_{13}a_{21} - a_{11}a_{23} & a_{11}a_{22} - a_{12}a_{21} \end{bmatrix}.$$

In the new coordinate system (λ') the coordinates $x_{i'}$ (i = 1, 2, 3, 3) 4, 5, 6) of the tensor $R_{\alpha\beta\gamma\delta}$ have the following form:

$$\begin{cases} x_{1'} = J^{-2} \left[a_{33}^2 x_1 + a_{23}^2 x_2 + a_{13}^2 x_3 - 2a_{23} a_{33} x_4 + 2a_{13} a_{33} x_5 - 2a_{13} a_{23} x_6 \right], \\ x_{2'} = J^{-2} \left[a_{32}^2 x_1 + a_{22}^2 x_2 + a_{12}^2 x_3 - 2a_{22} a_{32} x_4 + 2a_{12} a_{32} x_5 - 2a_{12} a_{22} x_6 \right], \\ x_{3'} = J^{-2} \left[a_{31}^2 x_1 + a_{21}^2 x_2 + a_{11}^2 x_3 - 2a_{21} a_{31} x_4 + 2a_{11} a_{31} x_5 - 2a_{11} a_{21} x_6 \right], \\ x_{4'} = J^{-2} \left[-a_{32} a_{33} x_1 - a_{22} a_{23} x_2 - a_{12} a_{13} x_3 + (a_{22} a_{33} + a_{23} a_{32}) x_4 - (a_{12} a_{33} + a_{13} a_{32}) x_5 + (a_{12} a_{23} + a_{13} a_{22}) x_6 \right], \\ x_{5'} = J^{-2} \left[a_{31} a_{33} x_1 + a_{21} a_{23} x_2 + a_{11} a_{13} x_3 - (a_{21} a_{33} + a_{23} a_{31}) x_4 + (a_{11} a_{33} + a_{13} a_{31}) x_5 - (a_{11} a_{23} + a_{13} a_{21}) x_6 \right], \\ x_{6'} = J^{-2} \left[-a_{31} a_{22} x_1 - a_{21} a_{22} x_2 - a_{11} a_{12} x_3 + (a_{21} a_{32} + a_{22} a_{31}) x_4 - (a_{11} a_{32} + a_{12} a_{31}) x_5 + (a_{11} a_{22} + a_{12} a_{21}) x_6 \right]. \end{cases}$$

Similarly to Section 2 the sought density H is a function of x_1, \ldots, x_6 fulfilling the functional equation:

(6)
$$H(x_1',\ldots,x_6')=\varepsilon|J|^rH(x_1,\ldots,x_6),$$

where

where
(7)
$$\varepsilon = \begin{cases} 1 & \text{for a Weyl density,} \\ \text{sgn} J & \text{for an ordinary density.} \end{cases}$$

Let us assume that $H(x_1, ..., x_6)$ is of class C^1 . For simplicity we introduce the following notation:

(8)
$$H_{i} \stackrel{\mathrm{df}}{=} \frac{\partial H}{\partial x_{i}}, \quad i = 1, 2, ..., 6.$$

First we differentiate the functional equation (6) with respect to the parameters a_{ij} , i, j = 1, 2, 3, and next we substitute

$$[a_{ij}] = [\delta_{ij}],$$

where δ_{ij} is the Kronecker delta.

re δ_{ij} is the Kronecker delta. Then we get a system of nine equations of the first order with one unknown function H depending on six variables x_i , i = 1, ..., 6:

$$\begin{cases} 2x_{1}H_{1} + 2x_{2}H_{2} + & +2x_{4}H_{4} + x_{5}H_{5} + x_{6}H_{6} = -rH, \\ 2x_{1}H_{1} + & +2x_{3}H_{3} + x_{4}H_{4} + 2x_{5}H_{5} + x_{6}H_{6} = -rH, \\ & 2x_{2}H_{2} + 2x_{3}H_{3} + x_{4}H_{4} + x_{5}H_{5} + 2x_{6}H_{6} = -rH, \\ & 2x_{6}H_{2} + & +x_{5}H_{4} + & +x_{3}H_{6} = 0, \\ 2x_{5}H_{1} + & +x_{6}H_{4} + x_{3}H_{5} & = 0, \\ 2x_{6}H_{3} + & +x_{4}H_{5} + x_{2}H_{6} = 0, \\ 2x_{4}H_{1} + & +x_{2}H_{4} + x_{6}H_{5} & = 0, \\ 2x_{5}H_{3} + & +x_{1}H_{5} + x_{4}H_{6} = 0, \\ 2x_{5}H_{3} + & +x_{1}H_{5} + x_{4}H_{6} = 0, \\ 2x_{4}H_{2} + & +x_{1}H_{4} + & +x_{5}H_{6} = 0. \end{cases}$$

We assert that (10) is a complete system. Denoting the left-hand sides of (10) by $X_1, X_2, ..., X_9$ respectively, we have

(11)
$$X_i(H) \stackrel{\text{df}}{=} \sum_{k=1}^6 a_{ik} H_k, \quad i = 1, 2, ..., 9,$$

where the coefficients a_{ik} are certain simple functions of x_1, \ldots, x_n . Let us introduce a shorter notation for Poisson brackets:

$$(12) (X_i, X_j) \stackrel{\text{df}}{=} \sum_{k=1}^{6} [X_i(a_{jk}) - X_j(a_{ik})] H_k, \quad i < j, i, j = 1, 2, ..., 9.$$

After a number of simple operations based on (10), (11) and (12) we get

$$(X_{1}, X_{2}) = 0, \qquad (X_{2}, X_{3}) = 0, \qquad (X_{3}, X_{5}) = X_{5},$$

$$(X_{1}, X_{3}) = 0, \qquad (X_{2}, X_{4}) = X_{4}, \qquad (X_{3}, X_{6}) = 0,$$

$$(X_{1}, X_{4}) = -X_{4}, \qquad (X_{2}, X_{5}) = 0, \qquad (X_{3}, X_{7}) = X_{7},$$

$$(X_{1}, X_{5}) = -X_{5}, \qquad (X_{2}, X_{6}) = 0, \qquad (X_{3}, X_{8}) = -X_{8},$$

$$(X_{1}, X_{6}) = X_{6}, \qquad (X_{2}, X_{7}) = X_{7}, \qquad (X_{3}, X_{8}) = -X_{8},$$

$$(X_{1}, X_{7}) = 0, \qquad (X_{2}, X_{8}) = 0, \qquad (X_{4}, X_{5}) = 0,$$

$$(X_{1}, X_{8}) = X_{8}, \qquad (X_{2}, X_{8}) = X_{9}, \qquad (X_{4}, X_{5}) = 0,$$

$$(X_{1}, X_{1}) = 0, \qquad (X_{2}, X_{1}) = X_{1}, \qquad (X_{2}, X_{2}) = X_{2},$$

$$(X_{1}, X_{2}) = 0, \qquad (X_{2}, X_{3}) = X_{3} - X_{1}, \qquad (X_{3}, X_{4}) = X_{5},$$

$$(X_{4}, X_{8}) = -X_{9}, \qquad (X_{5}, X_{8}) = X_{3} - X_{1}, \qquad (X_{6}, X_{9}) = -X_{8},$$

$$(X_{4}, X_{9}) = 0, \qquad (X_{5}, X_{9}) = X_{4}, \qquad (X_{7}, X_{8}) = X_{6},$$

$$(X_{5}, X_{6}) = -X_{7}, \qquad (X_{6}, X_{7}) = 0, \qquad (X_{8}, X_{9}) = 0,$$

from which it follows that (10) is a complete system.

For integrating complete systems of the type (10) it is convenient to find a so-called *integrating direction* ([15] or [10]).

Denoting the equations of the system (10) by (10.1)-(10.9), we shall integrate them in the following direction: (10.7), (10.5), (10.6), (10.8), (10.9), (10.4), (10.1), (10.2), (10.3).

To equation (10.7) corresponds the system of the ordinary equations:

(14)
$$\frac{dx_1}{2x_4} = \frac{dx_2}{0} = \frac{dx_3}{0} = \frac{dx_4}{x_2} = \frac{dx_5}{x_6} = \frac{dx_6}{0}.$$

Solving this system, we obtain

(15)
$$H = \varphi(x_2, x_3, x_6, x_4^2 - x_1 x_2, x_4 x_6 - x_2 x_5) = \varphi(y_1, \dots, y_5),$$
 where $\varphi \in C^1$.

Substituting solution (15) in equation (10.5) we get

(16)
$$\frac{dy_1}{0} = \frac{dy_2}{0} = \frac{dy_3}{0} = \frac{dy_4}{2y_5} = \frac{dy_5}{y_3^2 - y_1 y_2}.$$

Hence, by the assumption that $y_3^2 - y_1 y_2 \neq 0$, we have

(17)
$$H = \psi[x_2, x_3, x_6, x_2(x_1x_6^2 + x_2x_5^2 + x_3x_6^2 - x_1x_2x_3 - 2x_4x_5x_6)]$$
$$= \psi(z_1, z_2, z_3, z_4),$$

where $\psi \in C^1$.

Substituting solution (17) in equation (10.6) we get the system of equations

(18)
$$\frac{dz_1}{0} = \frac{dz_2}{2z_3} = \frac{dz_3}{z_1} = \frac{dz_4}{0}.$$

The solution of (18) is

(19)
$$H = \theta[x_2, x_6^2 - x_2x_3, x_2(x_1x_6^2 + x_2x_6^2 + x_3x_6^2 - x_1x_2x_3 - 2x_4x_5x_6)]$$
$$= \theta(u_1, u_2, u_3),$$

where $\theta \in C^1$.

Substituting solution (19) in equation (10.8) we get

$$(20) (x_4 x_6 - x_2 x_5) \theta_2 = 0,$$

hence, by assuming that $x_4x_6-x_2x_5\neq 0$, we receive

$$\theta_2 = 0$$
,

SO

(21)
$$H = \omega[x_2, x_2(x_1x_6^2 + x_2x_5^2 + x_3x_4^2 - x_1x_2x_3 - 2x_4x_5x_6)]$$
$$= \omega(v_1, v_2), \quad \omega \in C^1.$$

Substituting solution (21) in equation (10.9) we get

$$\frac{dv_1}{v_1} = \frac{dv_2}{v_2}.$$

The solution of (22) has the form

$$(23) H = \kappa(x_1x_2x_3 + 2x_4x_5x_6 - x_2x_5^2 - x_1x_6^2 - x_3x_4^2) = \kappa(w),$$

where $\varkappa(w) \in C^1$.

Solution (23) can be rewritten in the form

(24)
$$H = \varkappa(w) = \varkappa \begin{pmatrix} \begin{vmatrix} x_1 & x_4 & x_5 \\ x_4 & x_2 & x_6 \\ x_5 & x_6 & x_3 \end{vmatrix} \end{pmatrix}.$$

Substituting solution (23) in equation (10.4) we obtain an identity,

and thus equation (10.4) is not independent of the proviously integrated equations.

Substituting solution (24) in (10.1) we have

$$(25) 4x'(w)w = -rx(w).$$

Solving the homogeneous equation (25) we obtain (similarly to the equation (21), Section 2)

(26)
$$H = C|w|^{-\tau/4}, \quad r - \text{arbitrary}, \ w \neq 0,$$

where the integration constant $C \neq 0$ has the form

(27)
$$C = \begin{cases} C_1 & \text{for } w > 0, \\ C_2 & \text{for } w < 0, \end{cases}$$

and

(28)
$$w = \begin{vmatrix} x_1 & x_4 & x_5 \\ x_4 & x_2 & x_6 \\ x_5 & x_6 & x_3 \end{vmatrix}.$$

We verify without difficulty that solution (25) fulfils equations (10.2) and (10.3).

We also verify that the symmetric determinant (28), formed from the essential components (1) of the tensor $R_{a\beta\gamma\delta}$, is a Weyl density of weight (4), i.e. denoting by w' the right-hand side of (28) for $x_{i'}$, $i=1,2,\ldots,6$ of the form (5) we obtain (after tedious calculations)

(29)
$$w' = J^{-4}w$$
.

The results of Section 3 can now be formulated as follows:

THEOREM 2. In the space V_3 each scalar concomitant $H(x_1, ..., x_6)$ (of class C^1) of the curvature tensor $R_{\alpha\beta\gamma\delta}$ is a sonstant function $H(x_1, ..., x_6)$ = C, while every density concomitant of weight (-r) is a Weyl density of the form

(30)
$$H(x_1,\ldots,x_6)=C|w|^{-r/4}, \quad C\neq 0, \ w\neq 0,$$

where w is defined by (28) and C by (27).

Remark 1. The above considerations have only been based on the symmetry and antisymmetry of tensor $R_{a\beta\gamma\delta}$ but we have ignored the fact that $R_{a\beta\gamma\delta}$ as a curvature tensor comes from the metric tensor $g_{a\beta}$. The whole consideration is maintained if we assume that the tensor $R_{a\beta\gamma\delta}$ has only properties (10) and is independent of the tensor $g_{a\beta}$, i.e. the assumption that the space is Riemannian is not necessary.

Remark 2. The concomitant defined by formula (28) being an algebraic concomitant of the tensor $R_{a\beta\nu\delta}$, it can be called a differential

concomitant of the second order of the metric tensor $g_{\alpha\beta}$ (because $R_{\alpha\beta\gamma\delta}$ are expressed by $g_{\alpha\beta}$, $\partial_{\gamma}g_{\alpha\beta}$, $\partial_{\gamma\delta}g_{\alpha\beta}$). However, there are other algebraic concomitants of the tensor g, which are densities. The simplest of those is $g = \det(g_{\alpha\beta})$, another one (for n = 3) is a symmetric determinant of the third order formed from the essential components of the tensor

(31)
$$G_{\alpha\beta\gamma\delta} \stackrel{\mathrm{df}}{=} 2g_{[\alpha|\gamma|}g_{\beta|\delta},$$

i.e. from the minors of the second order of $\det(g_{\alpha\beta})$. The tensor $G_{\alpha\beta\gamma\delta}$ of the form (31) is the so-called induced metric tensor of the bivector space V_n^2 , occurring in paper [8].

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